

Write your student number on every sheet of paper.

**DO NOT** write your name on any of the papers.

It's not enough to do the calculations, you should also explain what you do.

1. Explain the following concepts/answer the following questions in a few lines. (Equations are not necessary, but if you use them, explain what the symbols are and what they mean.)
  - a) How does special relativity unify spatial rotations and Galilei transformations?
  - b) Explain the weak and strong equivalence principle **starting from general relativity**.
  - c) If we integrate the components of a tensor over a proper volume element, why do the resulting quantities **not** form the components of a tensor? (Excluding the case of scalars.)
2. Consider 1+1-dimensional Minkowski space in Cartesian coordinates.
  - a) Derive the Lorentz transformation matrix  $\Lambda^\alpha_\beta$  from the condition that the transformation leaves the line element

$$ds^2 = -dt^2 + dx^2$$

invariant. (You may neglect time reversal and spatial reflection.)

b) Express the transformation parameter in terms of the velocity  $v \equiv \frac{dx}{dt}$ .

3. Consider the spatially flat FLRW universe with the metric

$$ds^2 = -dt^2 + a(t)^2 \delta_{ij} dx^i dx^j .$$

Calculate the components  $R_{0i0j}$  and  $R_{ijkl}$  of the Riemann tensor.

4. Consider the linearly perturbed Minkowski metric

$$ds^2 = -(1 + 2\phi)dt^2 + (1 - 2\phi)\delta_{ij}dx^i dx^j .$$

Show that in the weak field and slow motion limit, the geodesic equation for timelike geodesics reduces to the Newtonian equation of motion  $\frac{d^2 x^i}{dt^2} = -\delta^{ij} \partial_j \phi$ .

5. Show that in the Schwarzschild spacetime photons can have circular geodesics. Show that such orbits are unstable.